

TRANSIENTS WITH NENUFAR

Few inputs for
the discussion

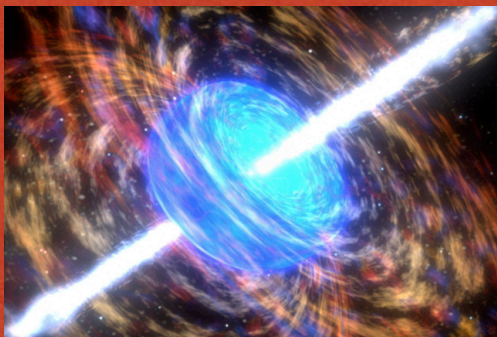
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14/2/2014



TWO FLAVOURS OF TRANSIENTS

- **Incoherent synchrotron emission**

- Relatively slow variability (\gg seconds)
- Associated with all explosive events
- Strong potential for MW astronomy



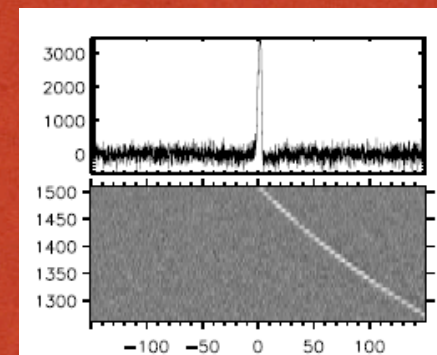
Detection: images

Expertise/Interests: Saclay, OCA?

- **Coherent emission**

- Relatively fast variability (\ll seconds)
- Usually associated with pulsars ?
- Often highly polarised

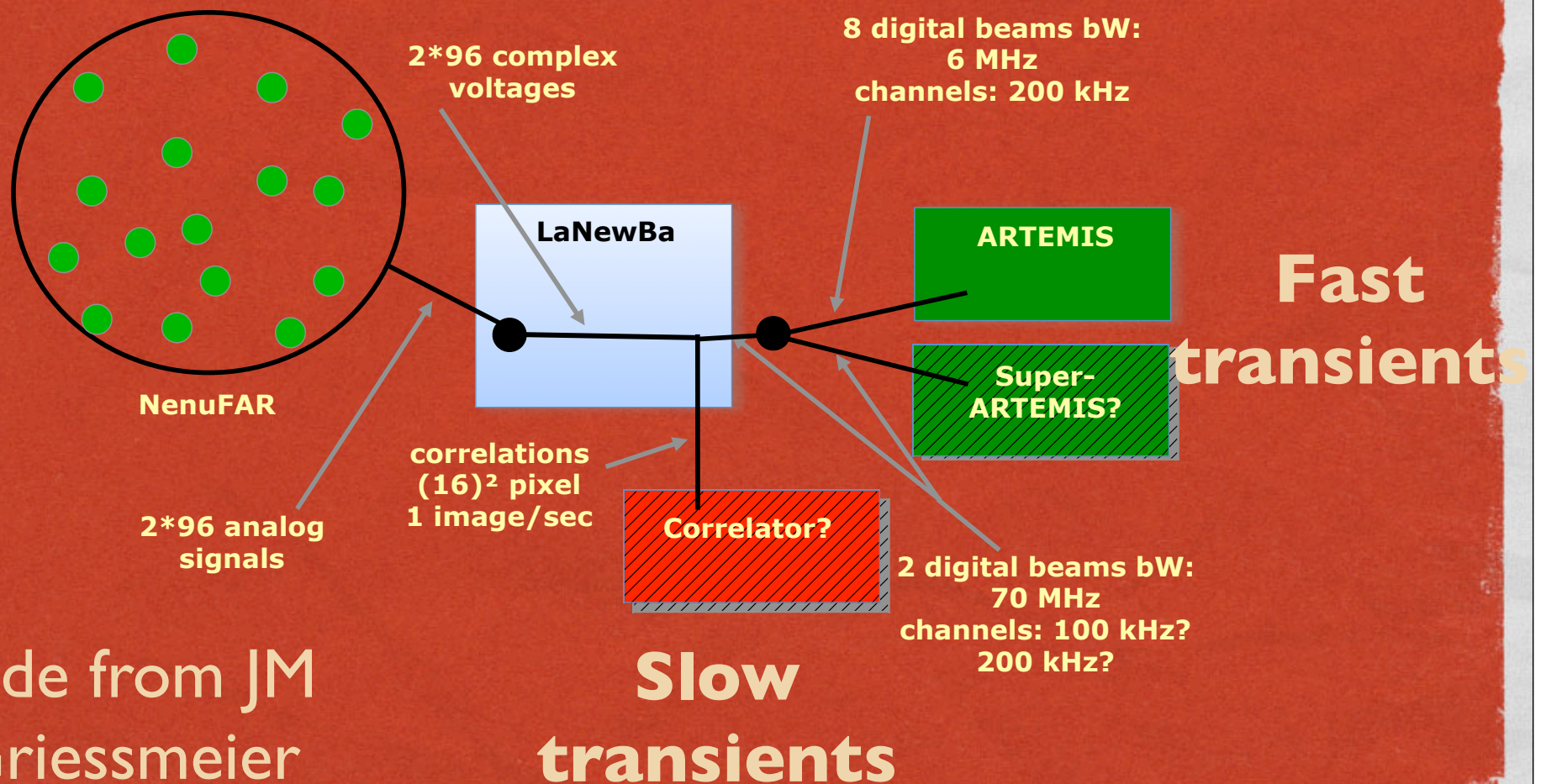
See SKA-LOFAR
radio days for
more details



Detection: time series

LPC2E, OP, Saclay?, ??

LANEWBA: A NEW DEDICATED BACKEND



Beamlets:

2 x 5 Gb/s

ARTEMIS-like backend

- Channelization
- pulsar dedispersion
- S-burst detection
- RFI mitigation ...

Subbands:

of MR x 2.5 Gb/s

NenuFAR-1: 75 Gb/s

NenuFAR: 480 Gb/s

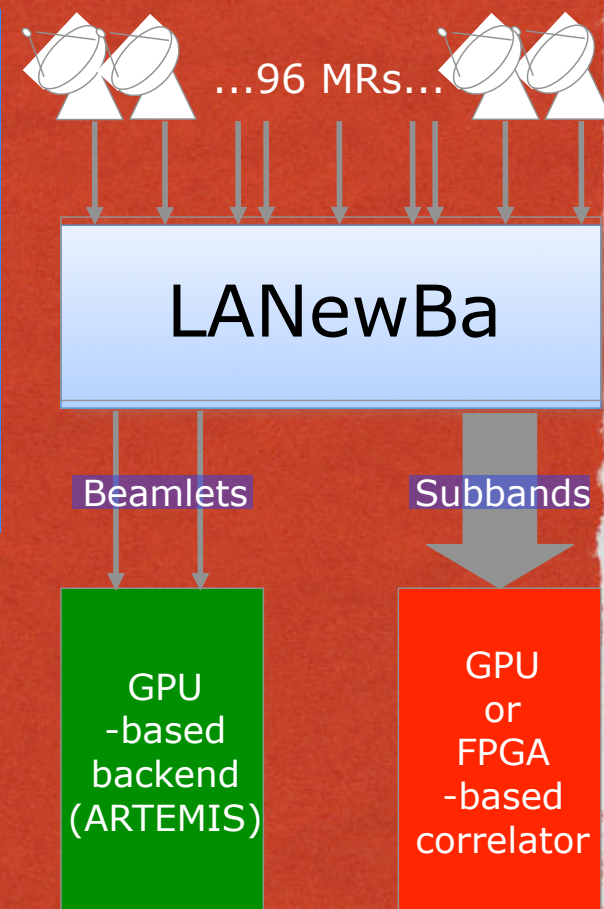
FPGA or GPU based correlator?

-> Spectro-polar-imaging

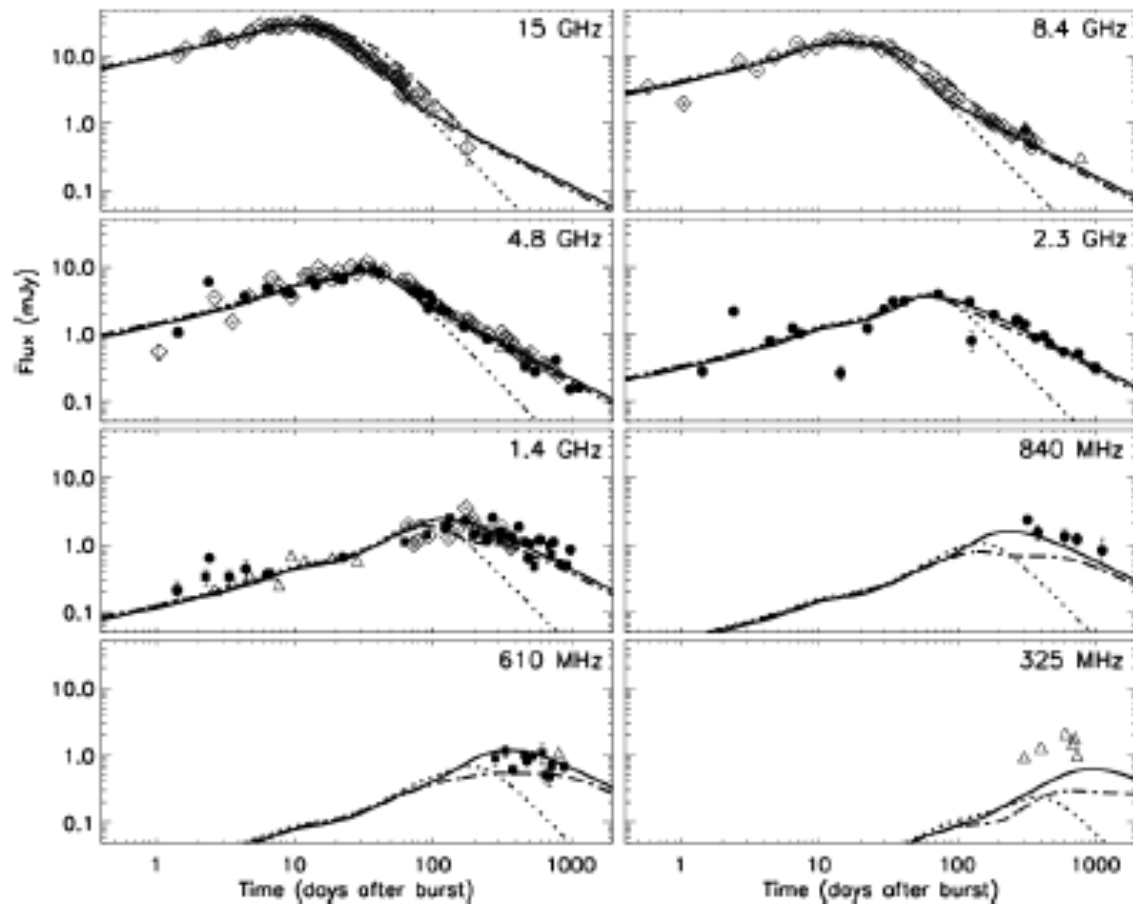
16x16-pixel skymap

~1s dump time

Transient search, Dark Ages, ...

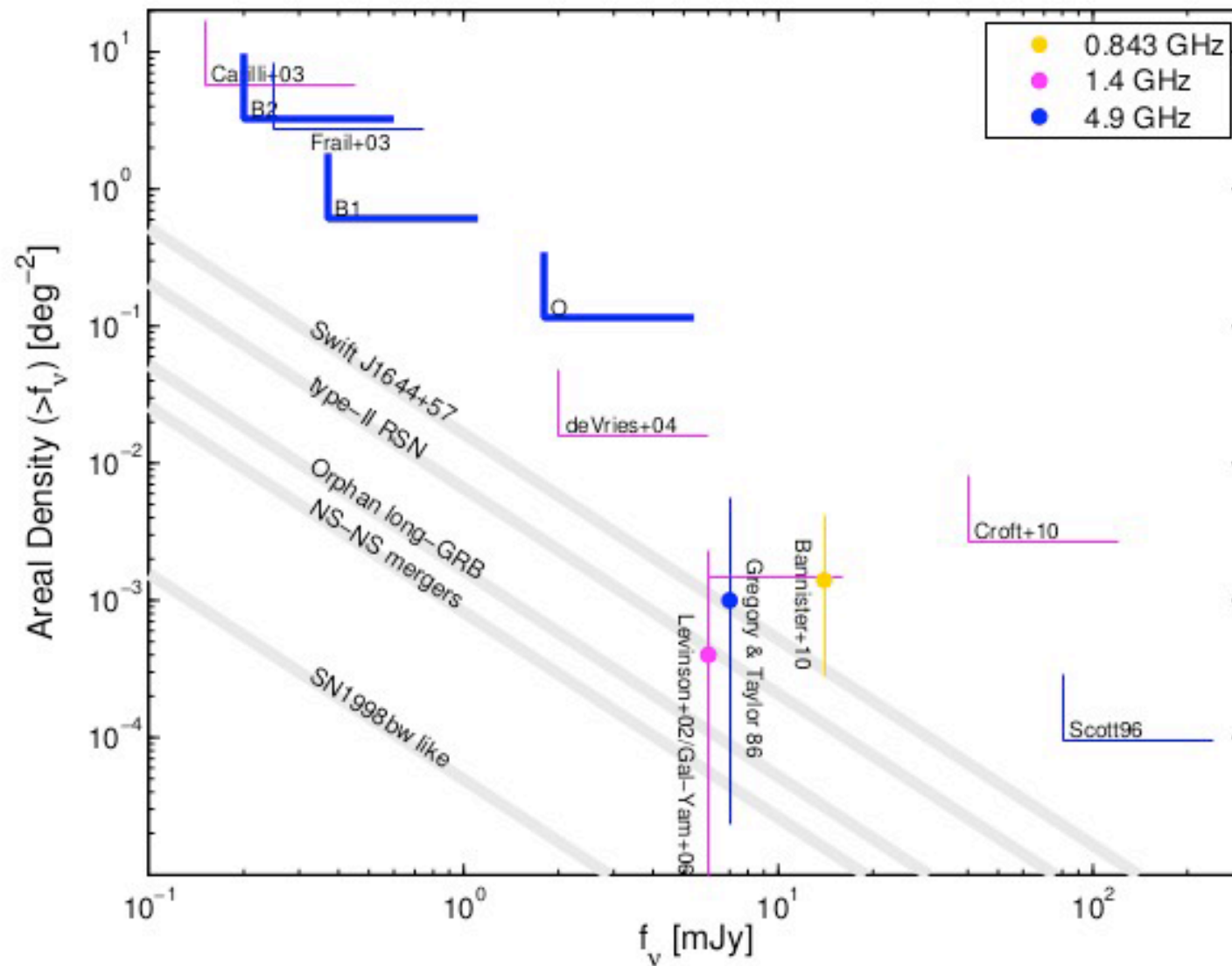


- Obviously, the two modes are potentially there for the two kind of transients.
- Is it really included in the design/cost definition ?
Transient buffer board ?
- Fast transients (see Ismael's talk). Probably the most interesting targets with higher impact
- Slow transients: capabilities to do fast imaging ? But we will likely be quite limited due to self-absorption

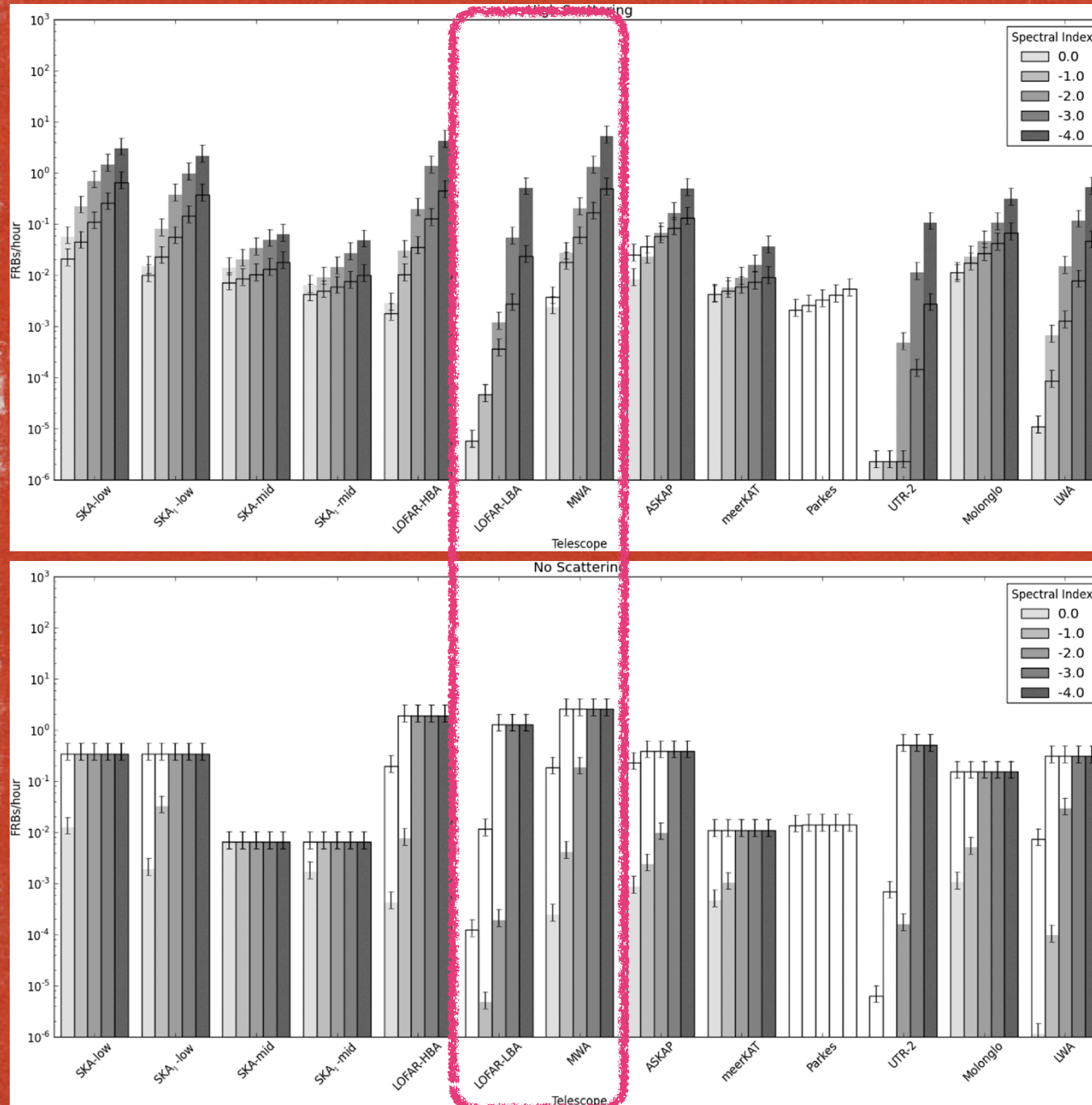


- Fast triggering possible ?
- Multi-stations triggering: Chilbolton ? removing false alert.

Radio transients



slow transients



- Prospect for FRB
- How does NenuFAR compare to these facility ?
- According to H13, in some cases, imaging survey more efficient than beamformed surveys.

CONCLUSIONS

- Good perspective, probably more on the fast transients side.
- Need to properly design/choose the back-end capabilities. Important to have these discussions now.
- Computing capabilities needed for fast transients ?
- «Not to Slow» transients or «slow fast» transient, capabilities for fast repointing : e.g. prompt emission from GRBs ? VO alert from GW.